Hydrodynamics of anisotropic quark and gluon fluids

Wojciech Florkowski, 1,2 Radoslaw Maj, 1 Radoslaw Ryblewski, 2 and Michael Strickland 3,4

1 Institute of Physics, Jan Kochanowski University, PL-25406 Kielce, Poland
2 The H. Niewodniczański Institute of Nuclear Physics,
Polish Academy of Sciences, PL-31342 Kraków, Poland
3 Physics Department, Gettysburg College, Gettysburg, PA 17325 United States
4 Frankfurt Institute for Advanced Studies, Ruth-Moufang-Strasse 1, D-60438, Frankfurt am Main, Germany
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The recently developed framework of anisotropic hydrodynamics is generalized to describe the dynamics of coupled quark and gluon fluids. The quark and gluon components of the fluids are characterized by different dynamical anisotropy parameters. The dynamical equations describing such mixtures are derived from kinetic theory with the collisional kernel treated in the relaxation-time approximation. Baryon number conservation is enforced in the quark and anti-quark components of the fluid, but overall parton number non-conservation is allowed in the system. The resulting equations are solved numerically in the (0+1)-dimensional boost-invariant case at zero and finite baryon density.

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I. INTRODUCTION

The dynamics of soft matter produced in ultrarelativistic heavy-ion collisions is well-described by relativistic viscous hydrodynamics [1–13]. However, due to the large gradients in the early stages of the collisions, viscous corrections to the ideal energy momentum tensor become quite substantial and the system becomes highly anisotropic in the momentum space, with the transverse pressure, P_{\perp} , typically much larger than the longitudinal pressure, P_{\parallel} . Large momentum-space anisotropies pose a problem for 2nd-order viscous hydrodynamics. since it relies on a linearization around an isotropic background. Large linear corrections generate unphysical results such as negative particle pressures, negative oneparticle distribution functions, etc. [14]. This has motivated the development of reorganizations of viscous hydrodynamics in which one incorporates the possibility of large momentum-space anisotropies at leading order. The framework developed has been dubbed anisotropic hydrodynamics [15–21].

In this paper we generalize the anisotropic hydrodynamics treatment to describe the dynamics of coupled quark and gluon fluids. Our approach closely follows the formulation presented in Refs. [16, 18, 20] which is based on kinetic theory with the collisional kernel treated in the relaxation time approximation. In this work the quark, antiquark, and gluon relaxation times are assumed to be constant and equal. We additionally assume that the system is boost invariant and homogeneous in the transverse directions which results in (0+1)-dimensional dynamics. As a result we can take the quark, antiquark, and gluon distribution functions to be of spheroidal form [22]; however, we allow for different dynamical anisotropy parameters in the quark and gluon sectors.

We begin by setting up the problem in terms of kinetic theory and then obtain the necessary dynamical equations by taking the zeroth and first moments of the

Boltzmann equation. We then require baryon number conservation in the quark/anti-quark sector. After doing this we find that the resulting system of equations is underdetermined. In order to resolve this problem we reduce the number of independent variables by requiring that the quark and anti-quark anisotropy parameters are equal.

The structure of the paper is as follows: In Sec. II we introduce the kinetic equations for quarks, antiquarks, and gluons in the relaxation time approximation. In Sec. III we discuss the zeroth moments of the kinetic equations. The baryon number and energy-momentum conservation laws are analyzed in Secs. IV and V, respectively. In Sec. VI the case of zero baryon density is discussed in greater detail. The results of our numerical calculations are presented in Sec. VII and VIII, for the cases with zero and non-zero baryon density, respectively. We conclude in Sec. IX. We close with an appendix in which we analyze the late-time behavior of the dynamical equations analytically.

II. KINETIC EQUATIONS

In this work, we consider kinetic equations for the phase space distribution functions of quarks and antiquarks, $Q^{\pm}(x, p)$, and gluons, G(x, p), given in the relaxation time approximation

$$p^{\mu}\partial_{\mu}Q^{\pm}(x,p) = -p^{\mu}U_{\mu}\frac{Q^{\pm}(x,p) - Q_{\text{eq}}^{\pm}(x,p)}{\tau_{\text{eq}}},$$
 (1)

$$p^{\mu}\partial_{\mu}G(x,p) = -p^{\mu}U_{\mu}\frac{G(x,p) - G_{\text{eq}}(x,p)}{\tau_{\text{eq}}}.$$
 (2)

Here $\tau_{\rm eq}$ is the relaxation time, and U^{μ} is the flow four-vector

$$U^{\mu} = \gamma(1, v_x, v_y, v_z), \gamma = (1 - v^2)^{-1}.$$
 (3)

In this paper we will consider a system that is transversely homogenous and undergoing only boost-invariant longitudinal expansion. In this case, the form of U^{μ} is fixed by the symmetry, and hence, below we set $v_x = v_y = 0$ and $v_z = z/t$.

Herein we will assume that the distribution functions are given by the covariant version of the Romatschke-Strickland distribution with f_{iso} given by a Boltzmann distribution [22]

$$Q^{\pm}(x,p) = \exp\left(\frac{\pm\lambda - \sqrt{(p \cdot U)^2 + \xi_q^{\pm}(p \cdot V)^2}}{\Lambda}\right),$$

$$G(x,p) = \exp\left(-\frac{\sqrt{(p \cdot U)^2 + \xi_g(p \cdot V)^2}}{\Lambda}\right), \quad (4)$$

where the transverse-momentum scale Λ is common for all of the distributions, while the anisotropy parameters ξ_q^{\pm} and ξ_g are allowed to be different. The parameter λ plays a role of the baryon chemical potential.

The corresponding equilibrium distributions $Q_{\text{eq}}^{\pm}(x, p)$ and $G_{\text{eq}}(x, p)$ are defined by the expressions

$$Q_{\text{eq}}^{\pm}(x,p) = \exp\left(\frac{\pm\mu - p \cdot U}{T}\right),$$

$$G_{\text{eq}}(x,p) = \exp\left(-\frac{p \cdot U}{T}\right),$$
(5)

where T is the temperature of the background and μ is the baryon chemical potential of quarks. The values of T and μ follow from the Landau matching conditions which originate from the energy-momentum and baryon number conservation laws.

The four-vector V^{μ} appearing in (4) defines the direction of the beam (z-axis)

$$V^{\mu} = \gamma_z(v_z, 0, 0, 1), \quad \gamma_z = (1 - v_z^2)^{-1/2}.$$
 (6)

We note that the four-vectors U^{μ} and V^{μ} satisfy the normalization conditions

$$U^2 = 1, \quad V^2 = -1, \quad U \cdot V = 0.$$
 (7)

In the local rest frame of the fluid element, U^{μ} and V^{μ} have simple forms

$$U^{\mu} = (1, 0, 0, 0), \quad V^{\mu} = (0, 0, 0, 1).$$
 (8)

For the (0+1)-dimensional boost-invariant expansion considered in this paper, we may use

$$U^{\mu} = (\cosh \eta, 0, 0, \sinh \eta),$$

$$V^{\mu} = (\sinh \eta, 0, 0, \cosh \eta),$$
(9)

where η is the space-time rapidity

$$\eta = \frac{1}{2} \ln \frac{t+z}{t-z}.\tag{10}$$

III. ZEROTH MOMENTS OF THE KINETIC EQUATIONS

Integrating Eqs. (1) and (2) over three-momentum and including the internal degrees of freedom we obtain the three equations

$$\partial_{\mu} N_q^{\pm \mu} = \frac{U_{\mu} \left(N_{q,\text{eq}}^{\pm \mu} - N_q^{\pm \mu} \right)}{\tau_{\text{eq}}}, \tag{11}$$

$$\partial_{\mu} N_g^{\mu} = \frac{U_{\mu} \left(N_{g, \text{eq}}^{\mu} - N_g^{\mu} \right)}{\tau_{\text{eq}}}, \tag{12}$$

where $N_q^{\pm \mu}$ and N_g^{μ} are particle currents

$$N_q^{\pm \mu} = n_q^{\pm} U^{\mu}, \quad N_g^{\mu} = n_g U^{\mu},$$
 (13)

$$N_{q,\text{eq}}^{\pm \mu} = n_{q,\text{eq}}^{\pm} U^{\mu}, \quad N_{q,\text{eq}}^{\mu} = n_{q,\text{eq}} U^{\mu}.$$
 (14)

This leads to the equations for the densities

$$\partial_{\mu}(n_q^{\pm}U^{\mu}) = \frac{n_{q,\text{eq}}^{\pm} - n_q^{\pm}}{\tau_{\text{eq}}},$$
 (15)

$$\partial_{\mu}(n_g U^{\mu}) = \frac{n_{g,\text{eq}} - n_g}{\tau_{\text{eq}}},$$
 (16)

where

$$n_q^{\pm} = \frac{g_q}{\pi^2} \frac{e^{\pm \lambda/\Lambda} \Lambda^3}{\sqrt{1 + \xi_q^{\pm}}}, \quad n_{q,eq}^{\pm} = \frac{g_q}{\pi^2} e^{\pm \mu/T} T^3, \quad (17)$$

$$n_g = \frac{g_g}{\pi^2} \frac{\Lambda^3}{\sqrt{1+\xi_g}}, \quad n_{g,eq} = \frac{g_g}{\pi^2} T^3.$$
 (18)

The factors g_q and g_g in Eqs. (17) and (18) account for internal degrees of freedom connected with spin and color. One can notice that they cancel out in (15) and (16). Substituting (17) and (18) into (15) and (16), respectively, we find the first dynamical equations, namely

$$\frac{3}{\Lambda} \frac{d\Lambda}{d\tau} - \frac{1}{2(1 + \xi_q^{\pm})} \frac{d\xi_q^{\pm}}{d\tau} \pm \frac{d}{d\tau} \frac{\lambda}{\Lambda} + \frac{1}{\tau}$$

$$= \frac{1}{\tau_{eq}} \left(\frac{e^{\pm(\mu/T - \lambda/\Lambda)} T^3}{\Lambda^3} \sqrt{1 + \xi_q^{\pm}} - 1 \right), (19)$$

for quarks, and

$$\frac{3}{\Lambda} \frac{d\Lambda}{d\tau} - \frac{1}{2(1+\xi_g)} \frac{d\xi_g}{d\tau} + \frac{1}{\tau}$$

$$= \frac{1}{\tau_{\text{eq}}} \left(\frac{T^3}{\Lambda^3} \sqrt{1+\xi_g} - 1 \right), \quad (20)$$

for gluons.

Equations (19) and (20) are three equations for seven unknown functions of the proper time: ξ_q^+ , ξ_q^- , ξ_g , Λ , λ , T, and μ . The remaining four equations will be obtained from the baryon and energy-momentum conservation laws discussed in the next two Sections.

IV. BARYON NUMBER CONSERVATION

The difference of Eqs. (15) for quarks and antiquarks, multiplied by a factor of 1/3, yields

$$\partial_{\mu}(bU^{\mu}) = \frac{b_{\rm eq} - b}{\tau_{\rm eq}},\tag{21}$$

where b is the quark baryon density

$$b = \frac{1}{3} (n_q^+ - n_q^-)$$

$$= \frac{g_q \Lambda^3}{3\pi^2} \left(\frac{e^{\lambda/\Lambda}}{\sqrt{1 + \xi_q^+}} - \frac{e^{-\lambda/\Lambda}}{\sqrt{1 + \xi_q^-}} \right), \quad (22)$$

and $b_{\rm eq}$ is the baryon density of the equilibrium background

$$b_{\text{eq}} = \frac{1}{3} \left(n_{q,\text{eq}}^{+} - n_{q,\text{eq}}^{-} \right)$$
$$= \frac{g_{q} T^{3}}{3\pi^{2}} \left(e^{\mu/T} - e^{-\mu/T} \right). \tag{23}$$

By demanding that the baryon number is conserved,

$$\partial_{\mu}(bU^{\mu}) = 0, \tag{24}$$

we obtain the Landau matching condition for the baryon density

$$b = b_{\text{eq}}. (25)$$

Equation (25) implies that the right-hand-sides of Eqs. (22) and (23) are equal. This results in a constraint connecting different parameters of our model.

To proceed further, we note that for (0+1)-dimensional expansion, the conservation law (21) has the Bjorken-type solution

$$b(\tau) = b_i \frac{\tau_i}{\tau},\tag{26}$$

which may be identified with the right hand side of (22). Here τ_i is the initial proper time and b_i is the corresponding initial baryon number density.

One may check that the use of the baryon number conservation (25) implies that Eqs. (15) and (16) are no longer independent. This suggests that the anisotropic distributions (4) contain too many parameters to be determined within our scheme. In order to have the same numbers of independent equations and free parameters, from now on, we assume that the quark and antiquark anisotropies are the same,

$$\xi_q^+ = \xi_q^- = \xi_q. \tag{27}$$

As a consequence, we find

$$\frac{\lambda}{\Lambda} = \sinh^{-1}(D) = \ln\left[D + \sqrt{1 + D^2}\right], \quad (28)$$

where, in order to simply our notation, we have introduced the quantity

$$D = \frac{3\pi^2 b\sqrt{1 + \xi_q}}{2g_q \Lambda^3}.$$
 (29)

Similarly, we find

$$\frac{\mu}{T} = \sinh^{-1}\left(\frac{D}{\kappa_q}\right) = \ln\left[\frac{D}{\kappa_q} + \sqrt{1 + \frac{D^2}{\kappa_q^2}}\right], \quad (30)$$

where

$$\kappa_q = \frac{T^3 \sqrt{1 + \xi_q}}{\Lambda^3}. (31)$$

In an analogous way we define

$$\kappa_g = \frac{T^3 \sqrt{1 + \xi_g}}{\Lambda^3}. (32)$$

Using Eqs. (28) and (30) in the "quark" equation in (19) one finds

$$\left(1 - \frac{D}{\sqrt{1+D^2}}\right) \left[\frac{3}{\Lambda} \frac{d\Lambda}{d\tau} - \frac{1}{2(1+\xi_q)} \frac{d\xi_q}{d\tau} + \frac{1}{\tau}\right]$$

$$= \left(1 - \frac{D}{\sqrt{1+D^2}}\right) \frac{\left[\kappa_q \sqrt{1+D^2} \sqrt{1+\frac{D^2}{\kappa_q^2}} - D^2 - 1\right]}{\tau_{\text{eq}}}.$$
(33)

Using Eqs. (28) and (30) in the "antiquark" equation in (19) one finds the same equation except that the factor $1-D/\sqrt{1+D^2}$ is replaced by $1+D/\sqrt{1+D^2}$. Since the factors $1\pm D/\sqrt{1+D^2}$ may be canceled, Eqs. (19) can be reduced to a single equation

$$\frac{3}{\Lambda} \frac{d\Lambda}{d\tau} - \frac{1}{2(1+\xi_q)} \frac{d\xi_q}{d\tau} + \frac{1}{\tau}$$

$$= \frac{1}{\tau_{\text{eq}}} \left[\kappa_q \sqrt{1+D^2} \sqrt{1+\frac{D^2}{\kappa_q^2}} - D^2 - 1 \right].$$
(34)

V. ENERGY-MOMENTUM CONSERVATION LAW

In order to introduce the energy-momentum conservation law for our system, we first sum over the quark, antiquark, and gluon degrees of freedom using Eqs. (1) and (2), then we multiply this sum by p^{ν} , and finally integrate over three-momentum. In this way, we obtain

$$\partial_{\mu}T^{\mu\nu} = \frac{U_{\mu} \left(T_{\text{eq}}^{\mu\nu} - T^{\mu\nu}\right)}{\tau_{\text{eq}}},\tag{35}$$

where [15, 20]

$$T^{\mu\nu} = (\varepsilon + P_\perp) U^\mu U^\nu - P_\perp g^{\mu\nu} - (P_\perp - P_\parallel) V^\mu V^\nu \,, \eqno(36)$$

and

$$T_{\rm eq}^{\mu\nu} = (\varepsilon_{\rm eq} + P_{\rm eq})U^{\mu}U^{\nu} - P_{\rm eq}g^{\mu\nu}. \tag{37}$$

The energy density appearing in (36) includes contributions from quarks, antiquarks, and gluons,

$$\varepsilon = \frac{3g_q \Lambda^4}{\pi^2} \left[2 \cosh(\lambda/\Lambda) \mathcal{R}(\xi_q) + r \mathcal{R}(\xi_g) \right]. \tag{38}$$

Here, to simplify our notation we have introduced the ratio of the degeneracy factors of gluons and quarks,¹

$$r = \frac{g_g}{g_q}. (39)$$

The function $\mathcal{R}(\xi)$ appearing in (38) has the form [16]

$$\mathcal{R}(\xi) = \frac{1}{2(1+\xi)} \left[1 + \frac{(1+\xi)\arctan\sqrt{\xi}}{\sqrt{\xi}} \right]. \tag{40}$$

Similar expressions can be given for the transverse and longitudinal pressures. In our case, the important quantity is the (longitudinal) enthalpy, $\varepsilon + P_{\parallel}$, which can be written in the compact form

$$\varepsilon + P_{\parallel} = -\frac{6g_q \Lambda^4}{\pi^2} \left[2 \cosh(\lambda/\Lambda) (1 + \xi_q) \mathcal{R}'(\xi_q) + r(1 + \xi_q) \mathcal{R}'(\xi_q) \right]. \tag{41}$$

Here, the prime denotes a derivative of the function $\mathcal{R}(\xi)$ with respect to its argument. The transverse pressure may be calculated from Eqs. (38) and (41), and the trace of the energy-momentum tensor which is zero for massless particles. The latter condition gives $\varepsilon = 2P_{\perp} + P_{\parallel}$.

In analogy to Eq. (38), we obtain the formula for the energy density of the equilibrium background

$$\varepsilon_{\rm eq} = \frac{3g_q T^4}{\pi^2} \left[2\cosh(\mu/T) + r \right], \tag{42}$$

In order to conserve energy and momentum, the right-hand-side of (35) should vanish. Hence, we obtain the Landau matching condition for the energy density

$$\varepsilon = \varepsilon_{\rm eq}.$$
 (43)

This leads us to another constraint

$$T^{4} = \Lambda^{4} \frac{2 \cosh(\lambda/\Lambda) \mathcal{R}(\xi_{q}) + r \mathcal{R}(\xi_{g})}{2 \cosh(\mu/T) + r},$$
 (44)

which can be written in the form

$$\frac{T^4}{\Lambda^4} = \frac{2\sqrt{1+D^2}\,\mathcal{R}(\xi_q) + r\,\mathcal{R}(\xi_g)}{2\sqrt{1+(3\pi^2b/2g_qT^3)^2} + r}.$$
 (45)

For a given set of the values of τ , $\xi_q(\tau)$, $\xi_g(\tau)$, and $\Lambda(\tau)$, Eq. (45) may be used to calculate the temperature of the equilibrium background, $T(\tau)$.

If the condition (43) is satisfied, the left-hand-side of Eq. (35) yields the conservation law for energy and momentum

$$\partial_{\mu}T^{\mu\nu} = 0. \tag{46}$$

In our (0+1)-dimensional case, Eq. (46) is reduced to the expression

$$\frac{d\varepsilon}{d\tau} = -\frac{\varepsilon + P_{\parallel}}{\tau}.\tag{47}$$

Substituting Eqs. (38) and (41) in (47) one finds

$$\frac{d}{d\tau} \left[\Lambda^4 \left(2\sqrt{1 + D^2} \, \mathcal{R}(\xi_q) + r \mathcal{R}(\xi_g) \right) \right]$$

$$= \frac{2\Lambda^4}{\tau} \left[2\sqrt{1 + D^2} \left(1 + \xi_q \right) \mathcal{R}'(\xi_q) + r (1 + \xi_g) \mathcal{R}'(\xi_g) \right].$$
(48)

Since we have eliminated the dependence on λ and μ , Eqs. (20), (34), (45), and (48) represent four equations for four unknown functions: ξ_q , ξ_g , T, and Λ .

VI. ZERO BARYON DENSITY

We consider first the special case of vanishing baryon number. The condition b=0 used in Eqs. (28) and (30) leads to the conclusions

$$\lambda = 0, \quad \mu = 0. \tag{49}$$

Hence, we are left with the following four equations for four unknown functions $(\xi_q, \xi_g, T, \text{ and } \Lambda)$

$$\frac{3}{\Lambda} \frac{d\Lambda}{d\tau} - \frac{1}{2(1+\xi_q)} \frac{d\xi_q}{d\tau} + \frac{1}{\tau} = \frac{\kappa_q - 1}{\tau_{\text{eq}}},\tag{50}$$

$$\frac{3}{\Lambda} \frac{d\Lambda}{d\tau} - \frac{1}{2(1+\xi_g)} \frac{d\xi_g}{d\tau} + \frac{1}{\tau} = \frac{\kappa_g - 1}{\tau_{\text{eq}}},\tag{51}$$

$$\frac{T^4}{\Lambda^4} = \frac{2\mathcal{R}(\xi_q) + r\mathcal{R}(\xi_g)}{2+r},\tag{52}$$

$$\frac{d}{d\tau} \left[\Lambda^4 \left(2\mathcal{R}(\xi_q) + r\mathcal{R}(\xi_g) \right) \right]
= \frac{2\Lambda^4}{\tau} \left[2(1+\xi_q)\mathcal{R}'(\xi_q) + r(1+\xi_g)\mathcal{R}'(\xi_g) \right]. \tag{53}$$

The above system of equations can be reduced to two equations, if we calculate the ratio T/Λ from Eq. (52) and the derivative $d\Lambda/d\tau$ from Eq. (53) and substitute these two quantities into Eqs. (50) and (51). In this way we obtain,

$$\frac{d\xi_q}{d\tau} = 2(1+\xi_q) \left[\frac{1}{\tau} + \frac{p_q(\xi_q, \xi_g)}{\tau_{\text{eq}} \Delta(\xi_q, \xi_g)} \right]$$
 (54)

 $^{^1}$ In our numerical calculations we use the values $g_q=2\cdot 2\cdot 3=12$ and $g_g=2\cdot 8=16,$ hence r=4/3.

and

$$\frac{d\xi_g}{d\tau} = 2(1+\xi_g) \left[\frac{1}{\tau} + \frac{p_g(\xi_q, \xi_g)}{\tau_{\text{eq}} \Delta(\xi_q, \xi_g)} \right], \quad (55)$$

where

$$\Delta(\xi_q, \xi_g) = \frac{\pi^2 (P_{\perp} - P_{\parallel})}{3g_q \Lambda^4} = 4\mathcal{R}(\xi_q) + 2r\mathcal{R}(\xi_g) + 6(1 + \xi_q)\mathcal{R}'(\xi_q) + 3r(1 + \xi_g)\mathcal{R}'(\xi_g), \quad (56)$$

$$p_{q}(\xi_{q}, \xi_{g}) = (1 - \kappa_{q}) \left[4\mathcal{R}(\xi_{q}) + 2r\mathcal{R}(\xi_{g}) \right] + 3r(\kappa_{g} - \kappa_{q})(1 + \xi_{g})\mathcal{R}'(\xi_{g}), \quad (57)$$

$$p_g(\xi_q, \xi_g) = (1 - \kappa_g) \left[4\mathcal{R}(\xi_q) + 2r\mathcal{R}(\xi_g) \right] + 6(\kappa_q - \kappa_g)(1 + \xi_g)\mathcal{R}'(\xi_g).$$
 (58)

One may check that

$$p_q - p_g = \Delta(\kappa_g - \kappa_q). \tag{59}$$

A. Zero baryon density: the case of small anisotropies

In order to analyze the case of small anisotropies we rewrite our system of dynamical equations (54) and (55) in the form

$$\Delta(\xi_q, \xi_g) \left[\frac{d\xi_q}{2(1+\xi_q)d\tau} - \frac{1}{\tau} \right] = \frac{p_q(\xi_q, \xi_g)}{\tau_{\text{eq}}}, \quad (60)$$

$$\Delta(\xi_q, \xi_g) \left[\frac{d\xi_g}{2(1+\xi_q)d\tau} - \frac{1}{\tau} \right] = \frac{p_g(\xi_q, \xi_g)}{\tau_{\text{eq}}}. \quad (61)$$

We are going to include the terms up to second order in ξ_q and ξ_g on both sides of (60) and (61). Since, the expansion of Δ starts with the linear terms

$$\Delta = \frac{4}{15} \left(2\xi_q + r\xi_g \right) - \frac{12}{35} \left(2\xi_q^2 + r\xi_g^2 \right) + \dots$$
 (62)

we may neglect ξ_q and ξ_g in the terms $(1+\xi_q)$ and $(1+\xi_g)$. Instead of using (60) and (61) directly, we consider now their difference and the sum. For the difference we obtain

$$\Delta(\xi_q, \xi_g) \left[\frac{d\xi_q}{2 d\tau} - \frac{d\xi_g}{2 d\tau} - \frac{\kappa_g - \kappa_g}{\tau_{eq}} \right] = 0. \quad (63)$$

This equation is fulfilled up to the second order if the equation in the square brackets is fulfilled in the first order. This leads to the equation

$$\frac{d}{d\tau}(\xi_q - \xi_g) = -\frac{\xi_q - \xi_g}{\tau_{\text{eq}}},\tag{64}$$

which for a constant relaxation time has the simple exponential solution

$$d \equiv \xi_q - \xi_g = Ae^{-\frac{\tau}{\tau_{\text{eq}}}}.$$
 (65)

Another useful quantity to consider is the linear combination

$$\xi \equiv \frac{2\xi_q + r\xi_g}{2 + r}.\tag{66}$$

Adding Eqs. (60) and (61) multiplied by 2 and r, respectively, and keeping again all the terms up to second order gives

$$\frac{\Delta^{(1)}}{2} \frac{d\xi}{d\tau} - \frac{\Delta^{(1)} + \Delta^{(2)}}{\tau} = \frac{p^{(2)}}{\tau_{eq}},\tag{67}$$

where $p^{(2)}$ includes the second-order terms of the quantity

$$p = \frac{2p_q + rp_g}{(2+r)}. (68)$$

Similarly, $\Delta^{(1)}$ and $\Delta^{(2)}$ denote the linear and quadratic terms in the expansion of Δ . Straightforward calculations yield

$$\Delta^{(1)} = \frac{4(2+r)\xi}{15},\tag{69}$$

$$\Delta^{(2)} = -\frac{12}{35} \frac{(2+r)^2 \xi^2 + 2rd^2}{2+r},\tag{70}$$

and

$$p^{(2)} = \frac{3rd^2}{10(2+r)} - \frac{(2+r)\xi^2}{15}. (71)$$

We note that the linear terms are absent in the expansion of p.

In Eq. (67) we may drop the term $\Delta^{(2)}/\tau$, since $\Delta^{(1)}/\tau$ is the leading term in the whole expression. In this way, we obtain

$$\xi \left(\frac{d\xi}{d\tau} - \frac{2}{\tau} \right) = \frac{1}{2\tau_{\text{eq}}} \left[\frac{9rA^2 e^{-2\tau/\tau_{\text{eq}}}}{2(2+r)^2} - \xi^2 \right]. \tag{72}$$

In the following, we search for the solutions of Eq. (72) in the form $\xi(\tau) = z(\tau)e^{-\tau/\tau_{\rm eq}}$. It is also convenient to introduce a dimensionless parameter $s = \tau/\tau_{\rm eq}$. In this way we obtain

$$\frac{dz}{ds} = \frac{2e^s}{s} + \frac{a^2 + z^2}{2z},\tag{73}$$

where a is a constant defined by the relation

$$a^2 = \frac{9rA^2}{2(2+r)^2} \,. \tag{74}$$

B. Zero baryon density: stability discussion

Equation (73) implies that z(s) is a monotonically increasing function of s. We consider now the positive and negative initial values of z, denoted as $z_0 = z(s_0)$, where $s_0 = \tau_0/\tau_{\rm eq}$ defines the time where our small-anisotropy expansion may become valid.

In the first case, we assume that $z(s) \gg a$ for $s \geq s_0$ (if z_0 is small, it grows very fast due to a singularity a^2/z). In this case we can approximate (73) by the equation

$$\frac{dz}{ds} = \frac{2e^s}{s} + \frac{z}{2},\tag{75}$$

which has an analytic solution of the form

$$z(s) = Ce^{s/2} + 2e^{s/2}\text{Ei}(s/2),$$
 (76)

which is consistent with the initial assumption that $z(s) \gg a$. The corresponding form of ξ is

$$\xi(s) = Ce^{-s/2} + 2e^{-s/2}\text{Ei}(s/2).$$
 (77)

This gives us the asymptotic behavior

$$\xi(s) = \frac{4}{s} + \frac{8}{s^2} + \dots = \frac{4\tau_{\text{eq}}}{\tau} + \frac{8\tau_{\text{eq}}^2}{\tau^2} + \dots$$
 (78)

In the second case, where z_0 is negative, the asymptotic solution of (73) is

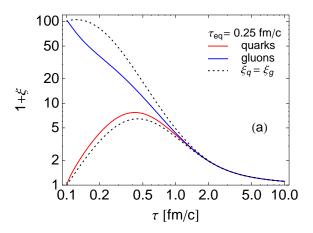
$$z(s) = -\frac{a^2 s}{4} e^{-s}. (79)$$

This implies an even faster decay of the anisotropy, since in this case

$$\xi(s) = -\frac{a^2s}{4}e^{-2s}. (80)$$

This, in turn, implies the very well satisfied approximate scaling $\xi_q = -(r/2)\xi_g$. As we see, in this case ξ_q and ξ_g will have opposite signs with both approaching zero. We note that when solving the equations numerically, one must take care since any small fluctuation of the negatively-valued ξ to positive values will cause the evolution the system of equations to approach the fixed point for the first case. Since both ξ 's approach zero exponentially the numerics will become sensitive to exponentially small round off errors. As a result one has to be careful with the numerics in this case.²

In App. A we present an alternative derivation of the late time behavior which does not rely on the change of variables used above.



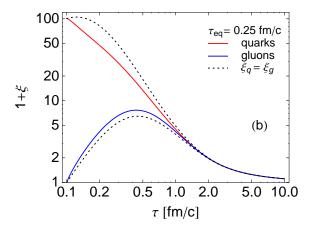


FIG. 1: (Color online) Time dependence of the anisotropy parameters of gluons (solid blue lines) and quarks (solid red lines) for the initial conditions $\xi_g=100$ and $\xi_q=0$ (a) and $\xi_q=100$ and $\xi_g=0$ (b).

VII. RESULTS FOR THE ZERO BARYON DENSITY

In this Section we discuss our results obtained in the case of vanishing baryon number density, b = 0. The solutions of Eqs. (54) and (55) are presented for a few possible options for the initial conditions.

In Fig. 1 we show the time dependence of the anisotropy parameters of gluons (solid blue lines) and quarks (solid red lines). The initial time for anisotropic hydrodynamic evolution is $\tau_i=0.1$ fm/c, and the relaxation time in the kinetic equations is constant and equals $\tau_{\rm eq}=0.25$ fm/c. In Fig. 1 (a), the initial values are $\xi_g^i=100$ and $\xi_q^i=0$, while in Fig. 1 (b) the initial values are reversed, $\xi_g^i=0$ and $\xi_q^i=100$. The dotted lines indicate the time evolution of the anisotropy parameters in the case where we have a one component system. In the latter case, the initial conditions correspond to either $\xi=100$ (upper dotted lines) or $\xi=0$ (lower dotted lines).

² We also note that we cannot use the initial conditions which give exactly $\Delta(\xi_q^i, \xi_g^i) = 0$. However, the probability that such initial conditions are realized in practice is negligible, since the measure of the curve $\Delta(\xi_q, \xi_g) = 0$ in the two-dimensional plane $\xi_q - \xi_g$ is zero.

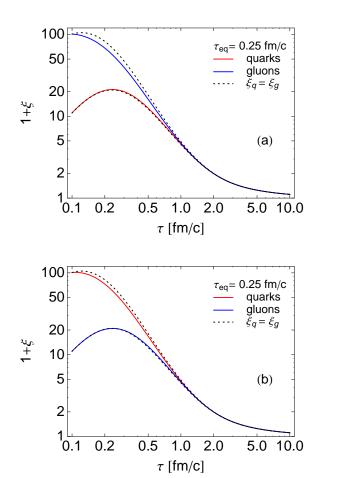


FIG. 2: (Color online) The same as Fig. 1 but for the initial conditions $\xi_g^i = 100$ and $\xi_q^i = 10$ (a) and $\xi_q^i = 100$ and $\xi_g^i = 10$ (b).

By comparing the solid and dotted lines in Fig. 1, we find that the anisotropy of a highly anisotropic subsystem interacting with a less anisotropic subsystem decreases faster, as compared to the case where a single highly anisotropic system is present. Similarly, the anisotropy of a less anisotropic subsystem is increased through the interaction with a more anisotropic subsystem. This behavior is naturally expected, on the grounds of the overall tendency of the whole system to reach equilibrium.

In Fig. 2 we show again the time dependence of the anisotropy parameters of gluons (solid blue lines) and quarks (solid red lines). In this case the initial conditions are $\xi_g^i=100$ and $\xi_q^i=10$ in (a) and $\xi_q^i=100$ and $\xi_g^i=10$ in (b). Since the differences between the quark and gluon anisotropies are smaller, than those presented in Fig. 1, the impact of the quark and gluon subsystems on each other is less significant — this can be inferred from the comparison of the solid and dotted lines.

The systems described in Figs. 1 and 2 correspond to systems which are initially *oblate* (the first anisotropy parameter is very large and positive, while the second

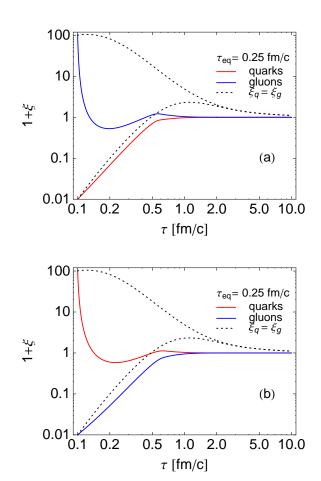
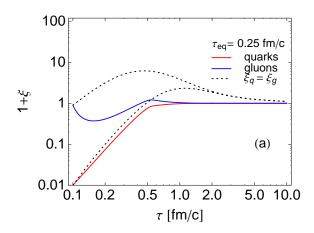


FIG. 3: (Color online) The time evolution of the quark and gluon anisotropies for the initially mixed prolate-oblate configurations: $\xi_g^i = 100$ and $\xi_q^i = -0.99$ (a) and $\xi_q^i = 100$ and $\xi_q^i = -0.99$ (b).

parameter is positive or equals zero). For comparison, it is interesting to consider also the cases where one or two anisotropy parameters are negative. Such cases correspond to mixed *oblate-prolate* or *prolate* configurations.

In Fig. 3 we show the results obtained for the initial conditions $\xi_g^i = 100$ and $\xi_q^i = -0.99$ (a) and $\xi_q^i = 100$ and $\xi_g^i = -0.99$ (b). In this case, we observe a very fast equilibration of matter. After about 0.5 fm/c, which corresponds to the evolution time $\tau \approx 2\tau_{\rm eq}$, the scaling solution $\xi_q = -(r/2)\xi_g$ is reached with an exponential decay of both anisotropies – see our discussion of the Case I in the appendix. Similar scaling behavior is observed in the cases where the two initial anisotropies are negative. This is shown in Fig. 4. We note that the systems which are initially oblate have the time asymptotics described by Case II in the appendix.



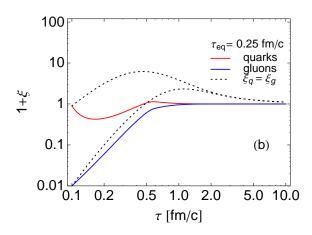


FIG. 4: (Color online) The time evolution of the quark and gluon anisotropies for the initial prolate configurations: $\xi_g^i = -0.1$ and $\xi_q^i = -0.99$ (a) and $\xi_q^i = -0.1$ and $\xi_g^i = -0.99$ (b).

VIII. RESULTS FOR THE FINITE BARYON DENSITY

The equations valid for b=0 may be also used to a good approximation in the case of small baryon number density, defined by the conditions

$$\frac{b}{T^3} \ll 1, \quad \frac{\mu}{T} \ll 1, \tag{81}$$

which, due to the Landau matching (25), imply also

$$\frac{b\sqrt{1+\xi_q}}{\Lambda^3} \ll 1, \quad \frac{\lambda}{\Lambda} \ll 1.$$
 (82)

This is so, because all corrections start with quadratic terms in b. In practice, the results obtained with the initial ratio $\lambda_i/\Lambda_i \sim 0.5$ are still very well approximated by the results obtained in the limit b=0. Therefore, to observe noticeable effects of baryon number conservation on the anisotropy evolution, we have to consider large initial values of the ratio λ/Λ .

In Fig. 5 we present our numerical results obtained for the initial ratio $\lambda_i/\Lambda_i=2$. We show the time dependence of the anisotropy parameters of gluons (solid blue lines) and quarks (solid red lines), with the same initial values as those used in Fig. 1, i.e., $\xi_g^i=100$ and $\xi_q^i=0$ in (a), and $\xi_g^i=0$ and $\xi_q^i=100$ in (b). The dashed lines describe the time dependence of the anisotropy parameters in the case b=0 (with the same initial conditions for anisotropy parameters). Compared to the case with b=0, we observe larger differences between the parts (a) and (b). Clearly, the baryon number conservation affects more strongly the evolution of quarks than the evolution of gluons — see the differences between the red and blue lines in the part (b). Figure 6 shows analogous results obtained with $\lambda_i/\Lambda_i=2$, $\xi_g^i=100$, and $\xi_q^i=10$ in (a), and $\xi_q^i=10$ and $\xi_q^i=10$ in (b).

We have also performed numerical calculations for the finite baryon density with the initial oblate-prolate and prolate configurations for $\lambda_i/\Lambda_i=2$. Our results are similar to those obtained at zero baryon density and presented in Figs. 3 and 4.

IX. CONCLUSIONS

In this paper we have generalized the concept of anisotropic hydrodynamics to describe the dynamics of coupled quark and gluon fluids. Our approach has followed closely the formulation presented in [16, 18, 20] which is based on kinetic theory with the collisional kernel treated in the relaxation time approximation. The resulting equations have been solved numerically in the (0+1)-dimensional boost-invariant case at zero and finite baryon density.

In our opinion, the results presented in this paper may be used to model the very early stages of ultra-relativistic heavy-ion collisions, where large pressure anisotropies are expected and the quark and gluon sectors may not have the same degree of momentum-space anisotropy. As a next step it would be interesting to generalize the approach presented here to a non-boost-invariant case, where the rapidity profiles of the baryon density and the energy density are different. This may be done along the lines presented in [17, 18]. It is also possible to relax the assumption that the isotropic distribution function is given by a Boltzmann distribution. This will not affect the dynamical equations qualitatively, and should only require the generalization of some of the dynamical equations presented herein. In addition, in the future one can also study the effect of including a relaxation time which is proportional to the inverse transverse momentum scale. We leave these interesting problems for future work.

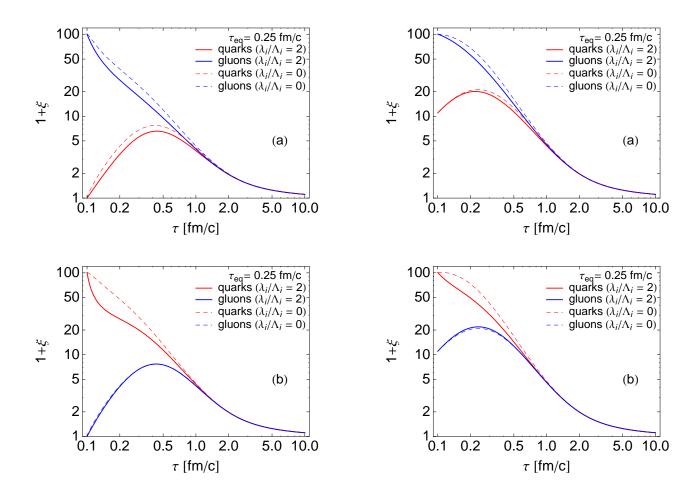


FIG. 5: (Color online) The same as Fig. 1 but with the inclusion of the finite baryon number density corresponding to $\lambda_i/\Lambda_i=2$. The dashed lines describe the time dependence of the anisotropy parameters in the case b=0.

FIG. 6: (Color online) The same as Fig. 2 but with the inclusion of the finite baryon number density corresponding to $\lambda_i/\Lambda_i = 2$. The dashed lines describe the time dependence of the anisotropy parameters in the case b = 0.

Acknowledgments

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Appendix A: Late-time behavior

In this appendix we present a discussion of the latetime behavior of the dynamical equations for the quark and gluon fluids. In order to gain a qualitative understanding of the dynamics one expects at late times, in Fig. 7 we present a vector plot of the right hand sides of Eqs. (54) and (55) when $\tau = 10^{12}$ fm/c. The vectors shown indicate the magnitude and phase-space direction of the time derivatives of ξ_q and ξ_g . The black line shows the line $\Delta(\xi_q, \xi_g) = 0$ with $\Delta(\xi_q, \xi_g)$ defined in Eq. (56). The red line shows the line $\xi_g = -(2/r)\xi_q$. The green line lindicates the line $\xi_g=\xi_q$. The darker shaded region corresponds to $\Delta(\xi_q,\xi_g)>0$ and the lighter shaded region to $\Delta(\xi_q,\xi_g)<0$.

As can be seen from this figure, the line $\Delta(\xi_q, \xi_g) = 0$ is a repulsive line, with points to the "right" of this line flowing to the right and points to the "left" of this line flowing to the left. In this way the system dynamically avoids the line of singularities corresponding to $\Delta(\xi_q, \xi_g) = 0$ (assuming that the system is not initialized with exactly $\Delta(\xi_q, \xi_g) = 0$).

Since at late times both anisotropy parameters tend towards zero, this analysis will help us to determine the late-time behavior of the dynamical equations. The final conclusions obtained are the same as those presented in

³ There is a subtle exception to this when $\xi_q = \xi_g$. In this case, it is possible to cross the line corresponding to $\Delta = 0$; however, there is no singularity in this case since the right hand sides of Eqs. (54) and (55) are finite at $\xi_q = \xi_g = 0$.

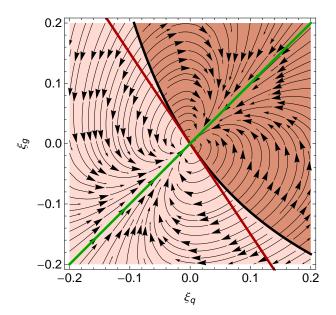


FIG. 7: (Color online) Phase space plot of the late time dynamics of the coupled dynamical equations for ξ_q and ξ_g . The black line shows the line $\Delta(\xi_q, \xi_g) = 0$. The red line shows the line $\xi_g = -(2/r)\xi_q$. The green line lindicates the line $\xi_g = \xi_q$. The darker shaded region corresponds to $\Delta(\xi_q, \xi_g) < 0$ and the lighter shaded region to $\Delta(\xi_q, \xi_g) > 0$.

Sec. VI A; however, the methods used are slightly different. Based on empirical analysis of numerical solutions to Eqs. (50), (51), and (53) there are two cases to be considered: (I) $\lim_{\tau\to\infty}\xi_q\neq\lim_{\tau\to\infty}\xi_g$ with $\lim_{\tau\to\infty}\xi_q/\xi_g<0$ and (II) $\lim_{\tau\to\infty}\xi_q=\lim_{\tau\to\infty}\xi_g=\xi>0$. Here we analyze the late-time solutions of the system in these two cases. To begin we take the difference of Eqs. (50) and (51) to obtain⁴

$$\frac{\dot{\xi}_q}{1+\xi_q} - \frac{\dot{\xi}_g}{1+\xi_g} = \frac{2(\kappa_g - \kappa_q)}{\tau_{\text{eq}}}.$$
 (A1)

where a dot indicates a derivative with respect to proper time. Since both ξ_q and ξ_g are small at late times, we can Taylor expand the equations in order to understand the late-time behavior. Using Eqs. (31), (32), and (52) we can expand to linear order in ξ_q and ξ_g to obtain

$$\dot{\xi}_q - \dot{\xi}_g = -\frac{\xi_q - \xi_g}{\tau_{\text{eq}}} \,, \tag{A2}$$

We can use this to solve for ξ_q in terms of ξ_g

$$\xi_q = \xi_g + Ae^{-s} \,, \tag{A3}$$

where A is an undetermined constant and $s \equiv \tau/\tau_{\rm eq}$. In case I one finds that ξ_g goes to zero exponentially in

 $\tau/\tau_{\rm eq}$ so one must keep the second term when solving for the asymptotic behavior. In case II, however, ξ_g goes to zero as an inverse power of $\tau/\tau_{\rm eq}$ so one can discard the second term by setting A=0.

a. Case I

We begin by using Eq. (53) to eliminate Λ' in Eqs. (50) and (51) and expand to linear order in ξ_q and ξ_g to obtain

$$\dot{\xi}_{g} - \dot{\xi}_{q} = \left(\frac{4}{15\tau} + \frac{1}{\tau_{eq}}\right) \xi_{q} + \left(\frac{2r}{15\tau} - \frac{1}{\tau_{eq}}\right) \xi_{g},
\dot{\xi}_{q} - \dot{\xi}_{g} = \left(\frac{4}{15\tau} + \frac{1}{\tau_{eq}}\right) \xi_{g} + \left(\frac{8}{15r\tau} - \frac{1}{\tau_{eq}}\right) \xi_{g}, (A4)$$

Adding these two equations we find $\xi_q = -r\xi_g/2$. Using (A3) we can then find the explicit forms for ξ_q and ξ_q

$$\lim_{\tau \to \infty} \xi_q = -\frac{1}{2} B r e^{-s},$$

$$\lim_{\tau \to \infty} \xi_g = B e^{-s},$$
(A5)

where B=-2A/(2+r) is an overall constant. Plugging these solutions back into Eq. (53) and taking the large- τ limit one obtains

$$\lim_{\tau \to \infty} \Lambda(\tau) = \frac{C}{\tau^{1/3}}, \tag{A6}$$

where C is an undetermined constant. These results describe the late-time behavior of the numerical solutions in Case I very well. As we can see from the expressions above, in this case one finds that at late times the system will, at some point, rapidly approach isotropy.

We note that it is possible to extend the solution above to next-to-leading order in the anisotropy parameters. To do this we expand ξ_q and ξ_q

$$\xi_q = -\frac{1}{2}Bre^{-s} + \delta\xi_q,$$

$$\xi_g = Be^{-s} + \delta\xi_g,$$
(A7)

treating the first terms as $\mathcal{O}(\xi_{q,g})$ and the second terms as $\mathcal{O}(\xi_{q,g}^2)$. Inserting these and expanding the resulting equations to quadratic order, one obtains the following solution for $\delta \xi_q$

$$\delta \xi_g = De^{-s} - \frac{a^2s}{4}e^{-2s}, \qquad (A8)$$

where D is an undetermined constant and $a^2 = 9rB^2/8$. Since the first term can be absorbed into the leading-order late-time behavior, we can set D = 0. With the resulting solution one finds

$$\delta \xi_g = \delta \xi_q = -\frac{a^2 s}{4} e^{-2s} \,, \tag{A9}$$

⁴ We could have also taken the difference of (51) over $1 + \xi_q$ and (55) over $1 + \xi_g$ and obtained the same equation.

Using this we see that the ξ defined via Eq. (66) simplifies to

$$\xi = -\frac{a^2s}{4}e^{-2s}\,,\tag{A10}$$

which is equivalent to the solution given by Eq. (80).

b. Case II

In this case the second term in (A3) can be discarded since empirically one finds that ξ_g goes to zero like an inverse power of τ . Substituting $\xi_q = \xi_g = \xi$ into either Eqs. (54) and (55) and expanding to linear order one obtains

$$\xi' = \frac{2}{\tau} + \left(\frac{2}{\tau} - \frac{1}{2\tau_{\rm eq}}\right)\xi. \tag{A11}$$

This equation can be solved analytically giving

$$\xi = \left[B + \operatorname{Ei}\left(\frac{\tau}{2\tau_{\text{eq}}}\right) \right] \frac{\tau^2}{4\tau_{\text{eq}}^2} e^{-\frac{\tau}{2\tau_{\text{eq}}}} - \frac{\tau}{2\tau_{\text{eq}}} - 1, \quad (A12)$$

where Ei is the exponential integral function and B is an undetermined constant. In the limit $\tau \to \infty$ one obtains

$$\lim_{\tau \to \infty} \xi(\tau) = \frac{4\tau_{\text{eq}}}{\tau} \,. \tag{A13}$$

Plugging this solution back into Eq. (53) and taking the large- τ limit one obtains

$$\lim_{\tau \to \infty} \Lambda(\tau) = \frac{C}{\tau^{1/3}}, \qquad (A14)$$

where C is an undetermined constant. Once again these results agree quite well with the numerical solutions in Case II.

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